SPICA Performance, DM, 26/01/2021

Numerical data and SNR models

1. Main hypothesis for the spectrograph

- a. R=140 for sensitivity and angular diameter measurements
 - i. Spectral band 600-900nm; 60 spectral channels of ~5nm (non-linear, from 3 to 7nm)
 - ii. for stars with θ =0.2mas, V²>0.6
 - iii. for stars with θ =0.5mas, V²>0.02
 - iv. for limb darkening (>0.8mas) V²_{min}=0.002
- b. R=4400 for spectral imaging
 - i. Spectral band of ~85nm in 600-900nm; 500 spectral channels of ~0.172nm
 - ii. Differential visibility measurements $V^2 = 0.6$ or $V^2 = 0.1$
- c. R=13100 for spectral imaging
 - i. Spectral band of ~29nm in 600-900nm; 500 spectral channels of ~0.057nm
 - ii. Differential visibility measurements $V^2 = 0.6$ or $V^2 = 0.1$

2. Detector parameters (Ixon897)

- a. Without fringe tracker: DIT=0.02s (group delay tracking only)
- b. Fringe tracker specification: DIT=0.2s (goal 10s with degraded transfer function)
- c. Quantum Efficiency qe=0.9, Read Out Noise RON=0.1
- d. 1 spectral channel=600 pixels
- e. N_{CIC}=0.0018 events/pixel
- f. N_{Dark}=0.00015 events/pixel/second

3. Transmission

- a. CHARA transmission: 0.017. Transmission AO=0.8.
- b. Pickup optics and injection, 4 mirrors: 0.994
- c. Fibres + spectrograph: 0.6; Splitter for photometry: 0.9
- d. Strehl=0.20; coupling efficiency=0.7xSR
- e. Collecting area: 6x0.75m²
- f. Instrumental visibility V=0.9

4. Model of Signal to Noise Ratio for V2 (Mourard et al., 2018 JOSA-A)

$$\text{a.} \quad SNR = \frac{\sqrt{N_{SpCh.N_{Fr}}} \left(\frac{N_{ph.V_{inst}.V_{target}}}{N_{Tel}}\right)^2}{\sqrt{PhotonNoise + ReadNoise + CoupledTerms}}.$$

b.
$$PhotonNoise = 2.\left(N_{ph} + N_{d+c}\right).\left(\frac{N_{ph}.V_{inst}.V_{target}}{N_{Tel}}\right)^2 + \left(N_{ph} + N_{d+c}\right)^2$$

c.
$$ReadNoise = N_{pix}.RON^2 + (N_{pix}.RON^2)^2$$

d. CoupledTerms =
$$2.N_{pix}.RON^2.\left(\frac{N_{ph}.V_{inst}.V_{target}}{N_{Tel}}\right)^2 + 2.\left(N_{ph} + N_{d+c}\right).N_{pix}.RON^2$$

- e. With the following notations
 - i. N_{SnCh} is the number of spectral channels considered
 - ii. N_{Fr} is the total number of frames
 - iii. N_{nh} is the number of photons per frame
 - iv. V_{inst} is the instrumental visibility
 - v. V_{target} is the visibility of the target
 - vi. N_{Tel} is the number of telescopes
 - vii. N_{d+c} is the number of dark and CIC events per frame
 - viii. RON is the Read Out Noise
 - ix. N_{nix} is the number of pixels considered

5. SNR on Differential Visibility

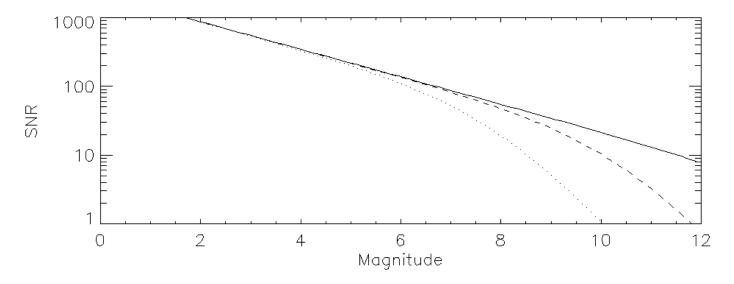
 Differential visibility is computed from the correlation of one wide spectral channel CH1 with a narrow spectral channel CH2.

- b. The SNR on Vdiff is computed as: $SNR_{Vdiff} = \sqrt{\frac{1}{\frac{1}{SNR_1^2} + \frac{1}{SNR_2^2}}}$, with SNR_1 and SNR_2 the SNR on V for the two channels. We assume $SNR_V = 2.SNR_{V^2}$.
- c. The uncertainty on the differential phase is given (in radian) by: $\sigma_{phi}=\frac{1}{SNR2}$

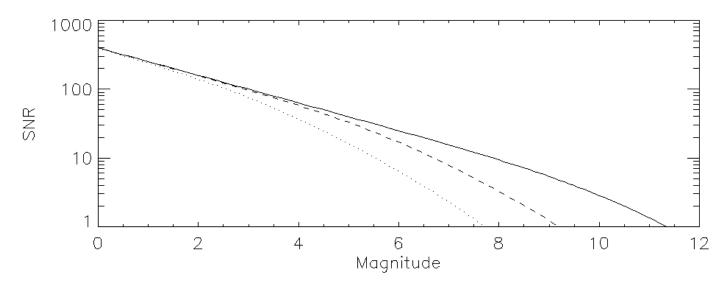
RESULTS on V² measurements

For all plots: dotted line: DIT=0.02s, dashed line: DIT=0.2s, solid line=DIT=10s

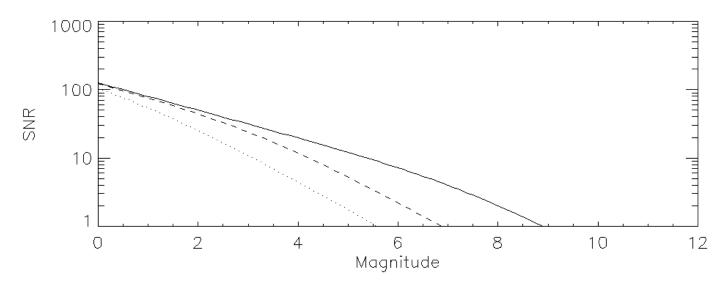
SNR on V² for one spectral channel: V²=0.6, 10mn of integration, R=140



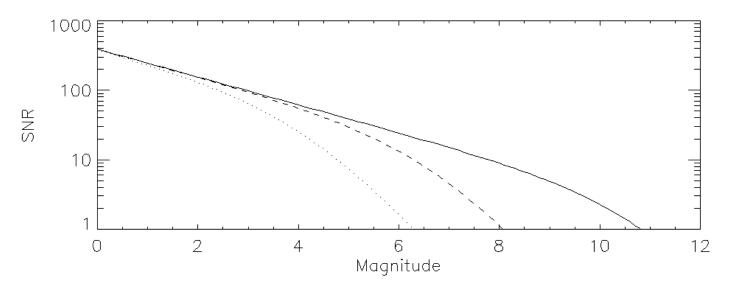
SNR on V² for one spectral channel: V²=0.02, 10mn of integration, R=140



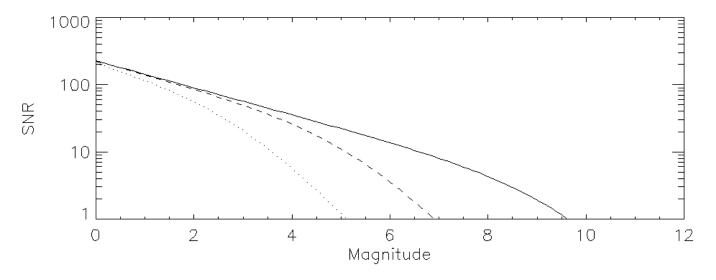
SNR on V² for one spectral channel: V²=0.002, 10mn of integration, R=140



SNR on V² for one spectral channel: V²=0.6, 10mn of integration, R=4400

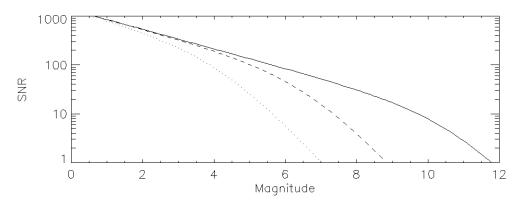


SNR on V² for one spectral channel: V²=0.6, 10mn of integration, R=13100

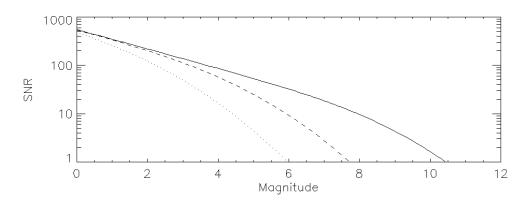


RESULTS on V diff measurements in medium and high-resolution mode (reference channel=250xscience channel)

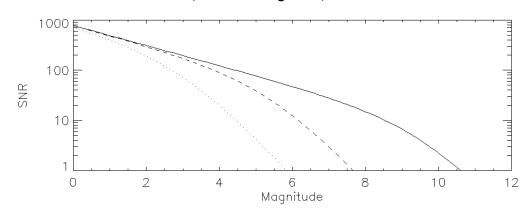
SNR on Vdiff, V²=0.6 in the reference channel, 30mn of integration, R=4400



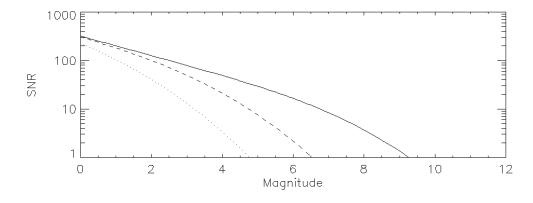
SNR on Vdiff, V²=0.1 in the reference channel, 30mn of integration, R=4400



SNR on Vdiff, V²=0.6 in the reference channel, 30mn of integration, R=13100

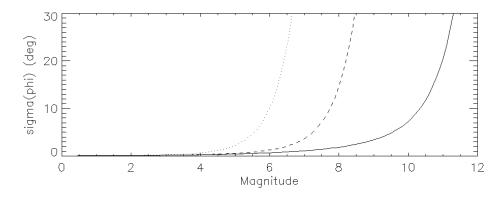


SNR on Vdiff, V²=0.1 in the reference channel, 30mn of integration, R=13100

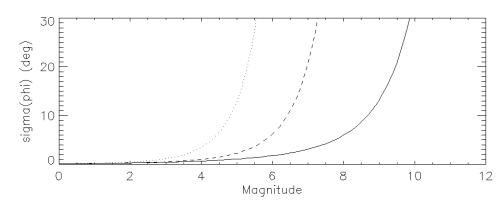


RESULTS on Differential Phase measurements in medium resolution mode

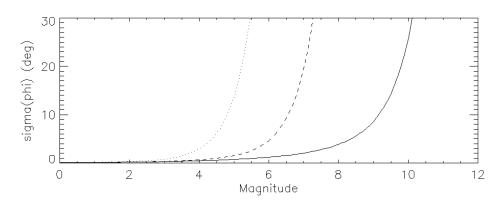
 σ_{phi} in deg, V²=0.6 in the reference channel, 30mn of integration, R=4400



 σ_{phi} in deg, V²=0.1 in the reference channel, 30mn of integration, R=4400



 σ_{phi} in deg, V²=0.6 in the reference channel, 30mn of integration, R=13100



 σ_{phi} in deg, V²=0.1 in the reference channel, 30mn of integration, R=13100

