

Tableau 1							
FRITE's NAME : FRITE MUSE - Formation, evolution and detection of galaxies							
Supervisor	D. Mary						
Tutors	E. Slezak, D. Mary						
Possible FRITE-elt(s). Indicate if repetition possible	F2, F3 or F5. Two occurrences at most.						
Required background	Physics : Quantum Mechanics and orbitals. Signal : maximum likelihood estimation.						
Theoretical/fondamental aspects	The FRITE contains two parts. The theoretical part provides a background in Cosmology with emphasis on galaxies' formation and evolution. The applications part focuses on how these galaxies are sensed and detected by the MUSE instrument.						
Related large instrument/project (if any)	MUSE integral field spectrograph @ VLT, Chile						
Keywords / Topics	<p>X Cosmology</p> <p>Turbulence</p> <p>Planetology</p> <p>Stellar Physics</p> <p>Relativity</p> <p>High Angular Resolution Instrumentation</p> <p>Numerical simulations</p> <p>X Signal/Image processing</p> <p>Mathematics</p> <p>OTHER :</p>						
Marks/evaluation mode	Theory : One final theoretical exam + 1 or 2 homeworks.						
	Applications : one final exam + numerical experiments on MUSE data.						
Documents/bibliography/on-line resources	http://muse.univ-lyon1.fr/soip.php?article99						
	R. Bacon et al., « The MUSE 3D view of the Hubble Deep Field South », A&A 575A, 2015, http://adsabs.harvard.edu/abs/2015A%26A...575A..75B						
FRITE's description	<p>* Theory : Formation and evolution of galaxies</p> <p>Tutor : E. Slezak</p> <p>Topics : Cosmology and Standard model. Quantum mechanics : Hydrogen lines. Classification of galaxies.</p> <p>Overview</p> <p>I. Basics in cosmology : the standard model and the evolution of the Universe</p> <p>II. Classification of galaxies (spirals, elliptics, active/non active).</p> <p>III. Dynamics in galaxies. Importance of observing young galaxies</p> <p>IV. Emission lines of Hydrogen. The Lyman series.</p> <p>V. Lyman alpha emitters and Lyman alpha forests. Quantum phenomena used to trace Universe on cosmological scales.</p> <p>VI. Formation and evolution galaxies as a major scientific goals of MUSE.</p> <p>Comparison of MUSE w.r.t. main existing surveys (HST,...).</p>						
	<p>* Application : Detection methods applied to Lyman alpha emitters for the MUSE instrument</p> <p>Tutor : D. Mary</p> <p>Topics : MUSE instrument. Statistical models. Detection theory.</p> <p>Overview</p> <p>I. Main characteristics of the MUSE instrument: suitability of the instrumental design for detecting distant galaxie. Observing strategy and main characteristics of the resulting data cube.</p> <p>Survey of past/current instruments addressing the same problem. Major related ground based facilities/organisms (ESO, VLT etc).</p> <p>Integral-field spectrographs in Astronomy.</p> <p>II. Introduction to statistical decision theory. Neyman-Pearson theorem, ROC curves, binary vs multiple hypothesis testing.</p> <p>III. Composite hypotheses. The GLRT.</p> <p>IV. Application to the detection of Lyman alpha emitters. Statistical models for MUSE data cubes.</p>						